

# The Neutron Lifetime Experiment $\tau$ SPECT

Dieter Ries  
dieter.ries@psi.ch

Workshop on UCN and VCN sources at the INP, Kazakhstan

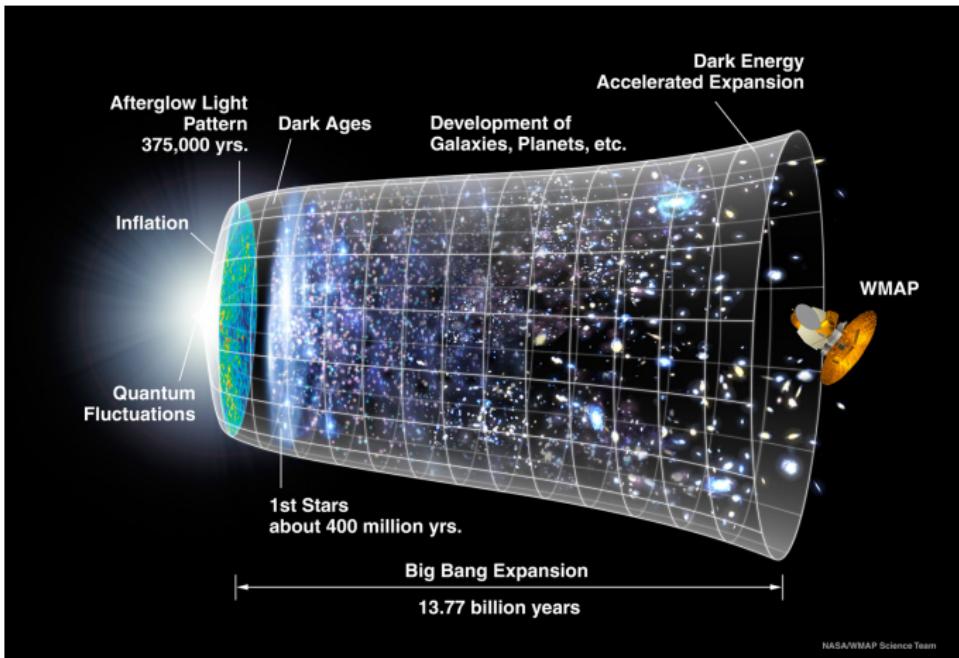
April 8, 2024

# Why Neutron Lifetime?

## a) Big Bang Nucleosynthesis (He abundance)

[Cyburt et al., doi:10.1103/RevModPhys.88, 2016]

# Big Bang Nucleosynthesis



@ $t = 2$  min:  $n/p \simeq 1/6$

@ $t = 4$  min:  $n/p \simeq 1/7$

# Neutron Lifetime

## Why n-lifetime?

- a) Big Bang Nucleosynthesis (He abundance)

[Cyburt et al., doi:10.1103/RevModPhys.88, 2016]

- b) CKM Unitarity ( $V_{ud}$ )

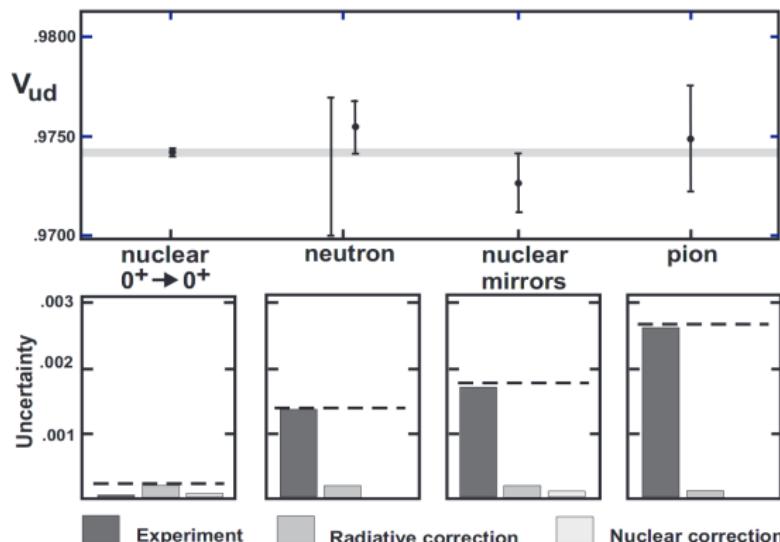
[Czarnecki, Marciano, Sirlin, doi:10.1103/PhysRevD.100.073008, 2019]

# Cabibbo–Kobayashi–Maskawa matrix

$$\begin{bmatrix} |V_{ud}| & |V_{us}| & |V_{ub}| \\ |V_{cd}| & |V_{cs}| & |V_{cb}| \\ |V_{td}| & |V_{ts}| & |V_{tb}| \end{bmatrix}$$

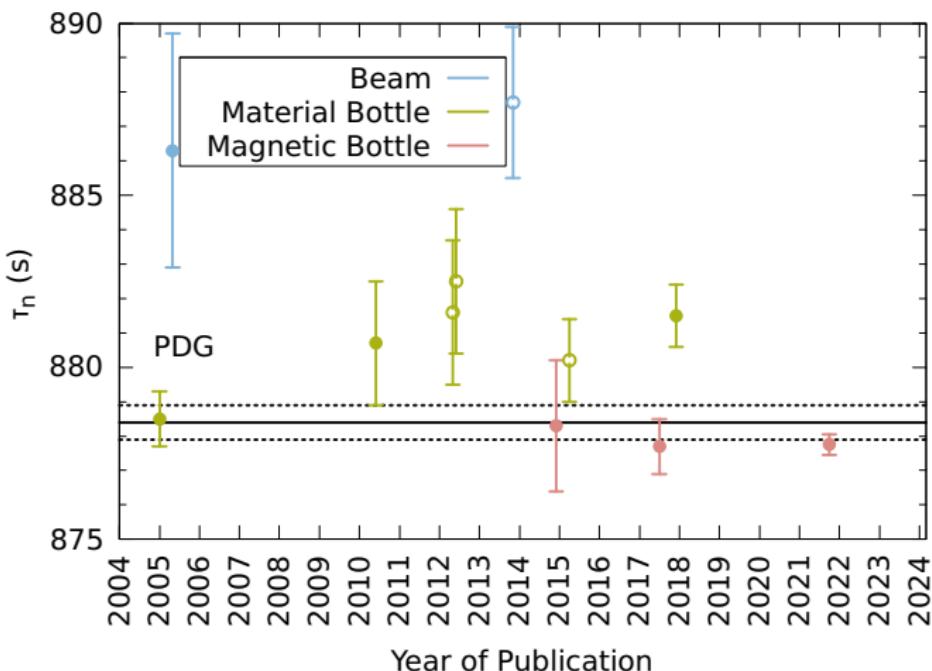
## Cabibbo–Kobayashi–Maskawa matrix

$$\begin{bmatrix} |V_{ud}| & |V_{us}| & |V_{ub}| \\ |V_{cd}| & |V_{cs}| & |V_{cb}| \\ |V_{td}| & |V_{ts}| & |V_{tb}| \end{bmatrix}$$



[Hardy and Towner, doi:10.48550/arXiv.1807.01146, 2018]

# The Lifetime Puzzle



# Neutron Lifetime

## Why n-lifetime?

- a) Big Bang Nucleosynthesis (He abundance)

[Cyburt et al., doi:10.1103/RevModPhys.88, 2016]

- b) CKM Unitarity ( $V_{ud}$ )

[Marciano and Sirlin, doi:10.1103/PhysRevLett.96.032002, 2006]

- c) “It’s 2024. We cannot agree on  $\tau_n$  to better than 10s?!”

$$\tau_{n,\text{beam}} = 887.7 \pm 1.2 \pm 1.9\text{s}$$

≠

$$\tau_{n,\text{stored}} = 877.75 \pm 0.28 \pm 0.22\text{s}$$

# $\tau$ SPECT

## Concept:

- 3-D magnetic storage
  - Two solenoids + Octupole

# $\tau$ SPECT

## Concept:

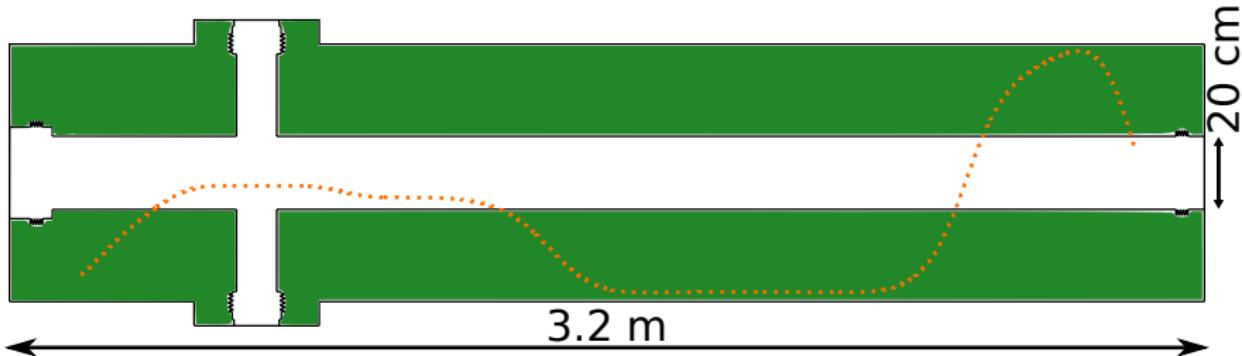
- 3-D magnetic storage
  - Two solenoids + Octupole
- Spinflip-loading
  - Holding field polarizes neutrons
  - Fast adiabatic spinflip as loading mechanism

# $\tau$ SPECT

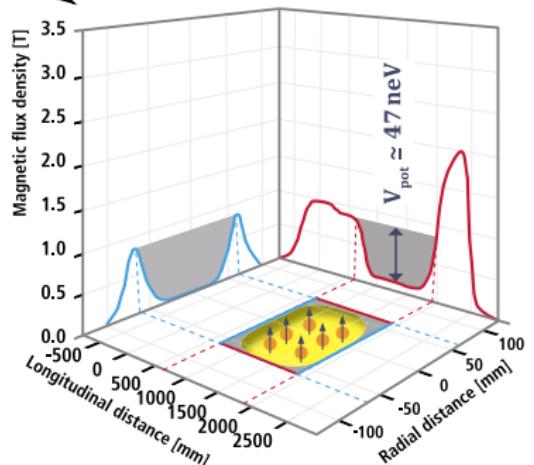
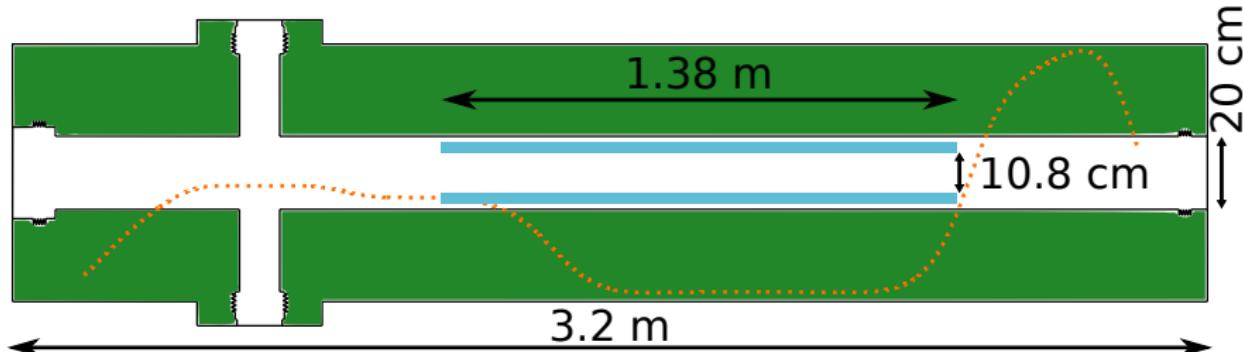
## Concept:

- 3-D magnetic storage
  - Two solenoids + Octupole
- Spinflip-loading
  - Holding field polarizes neutrons
  - Fast adiabatic spinflip as loading mechanism
- In-situ UCN detection
  - Minimizes extraction losses
  - High detector requirements wrt temp. & B-field

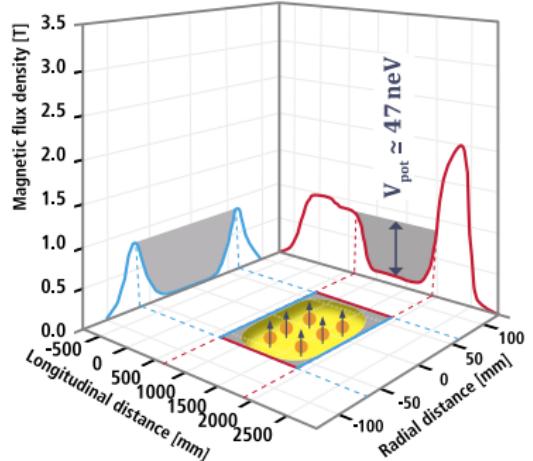
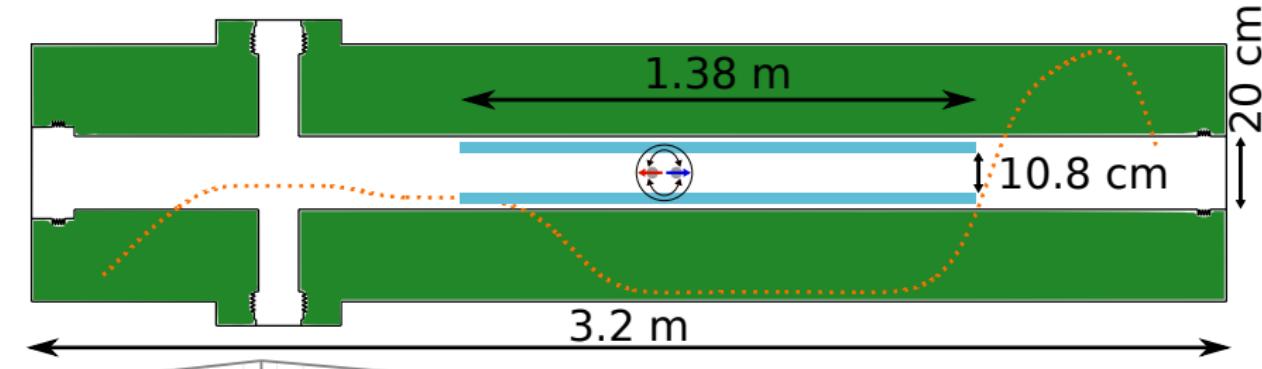
# $\tau$ SPECT fields



# $\tau$ SPECT fields

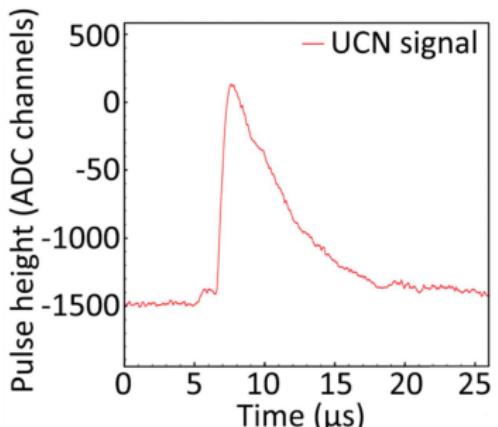
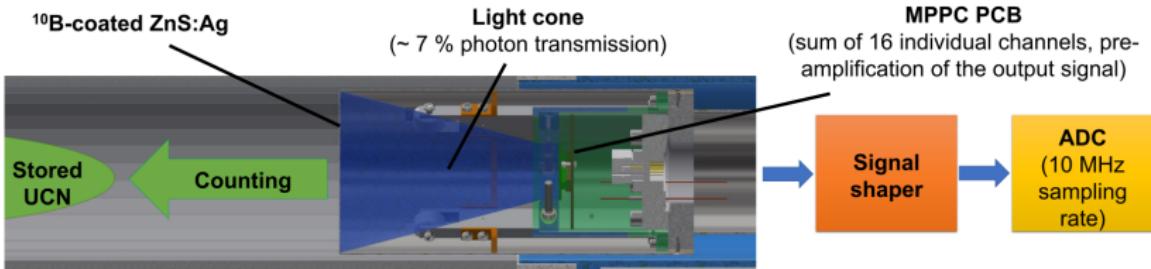


# $\tau$ SPECT fields



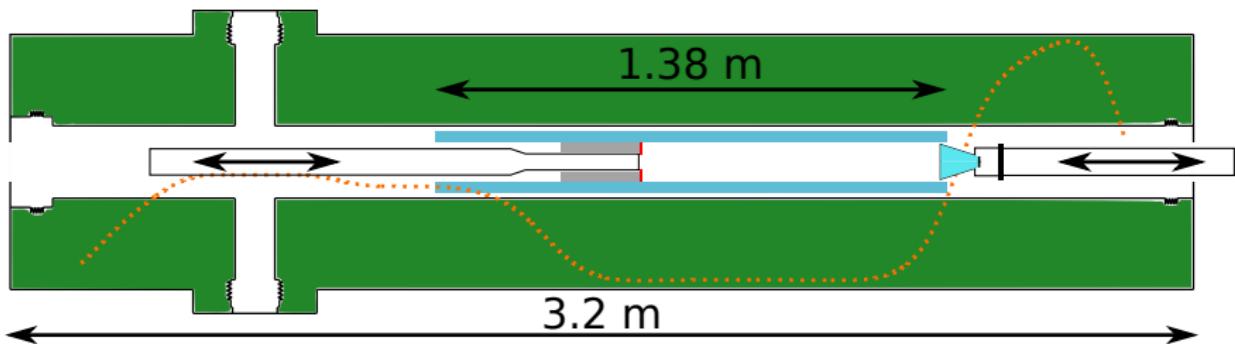
K. U. Ross

# $\tau$ SPECT Detector



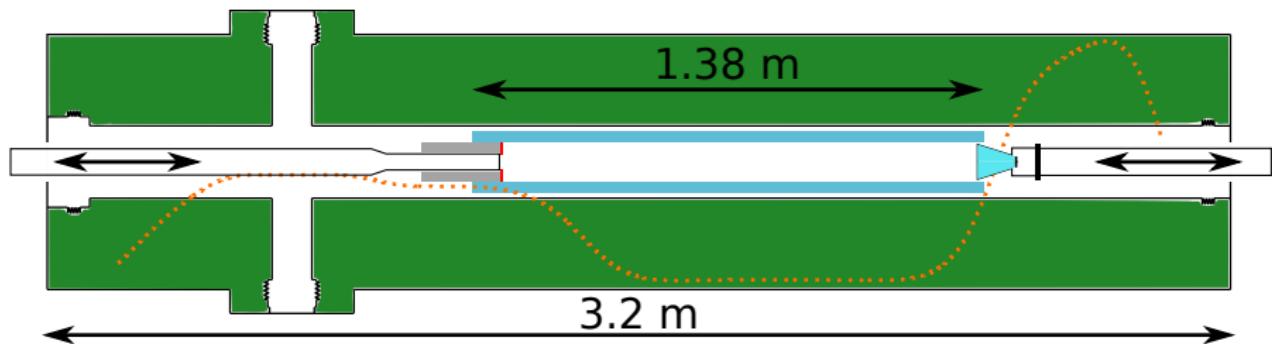
PhD work of J. Kahlenberg

# Measurement Procedure



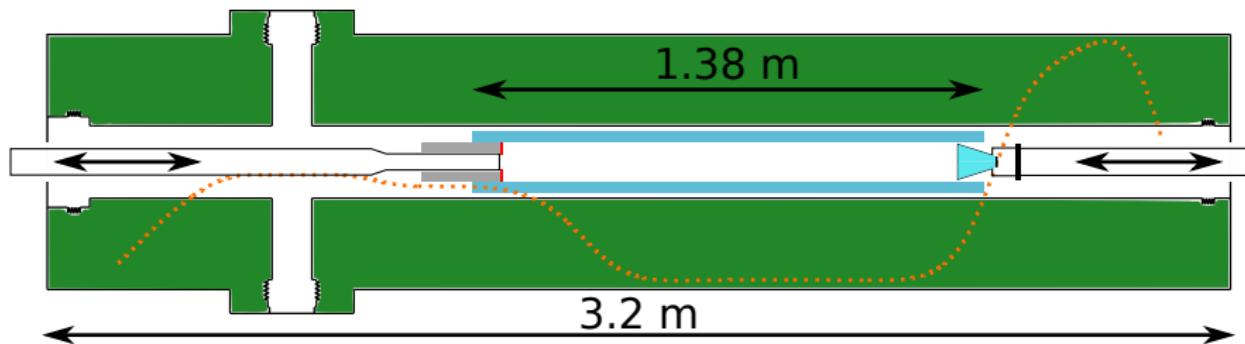
1. Fill UCN into  $\tau$ SPECT Magnet from the left
  - Polarization due to high Magnetic Field, SF on
  - Simultaneously: Intensity Monitoring (non-trappable UCN)

# Measurement Procedure



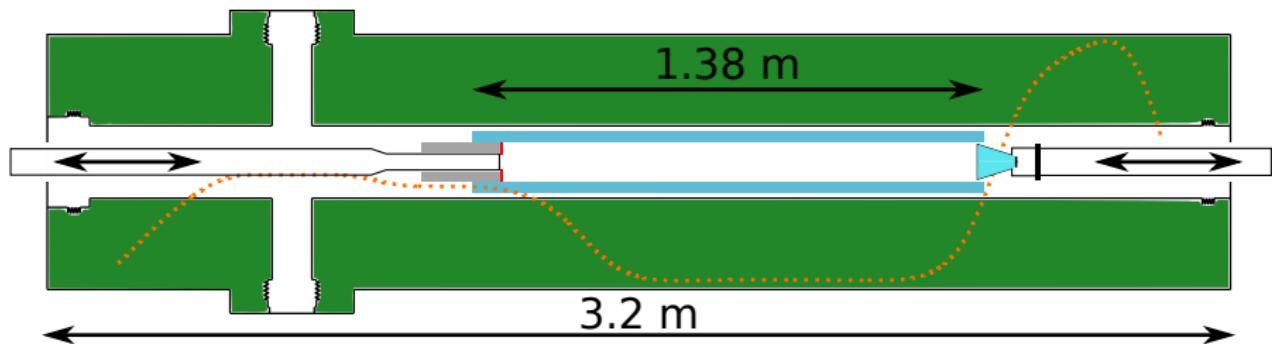
1. Fill UCN into  $\tau$ SPECT Magnet from the left
  - Polarization due to high Magnetic Field, SF on
  - Simultaneously: Intensity Monitoring (non-trappable UCN)
2. Remove SF from storage region

# Measurement Procedure



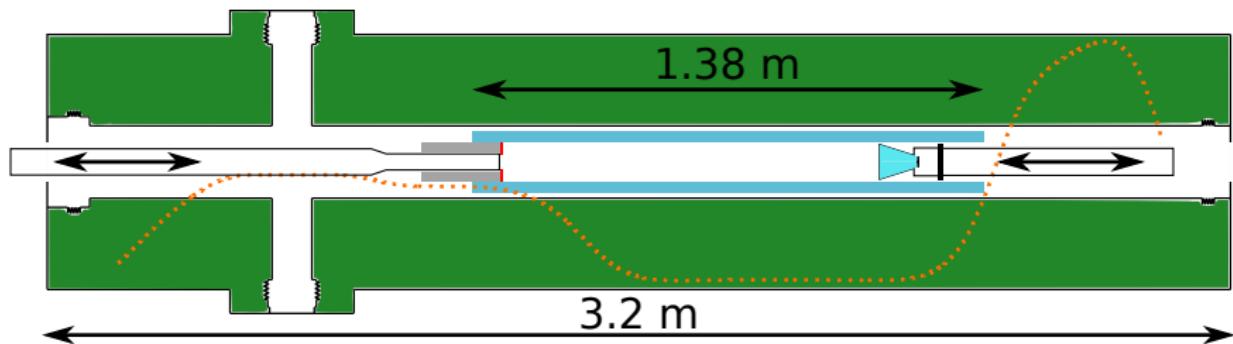
1. Fill UCN into  $\tau$ SPECT Magnet from the left
  - Polarization due to high Magnetic Field, SF on
  - Simultaneously: Intensity Monitoring (non-trappable UCN)
2. Remove SF from storage region
3. Detector to cleaning position and back

# Measurement Procedure



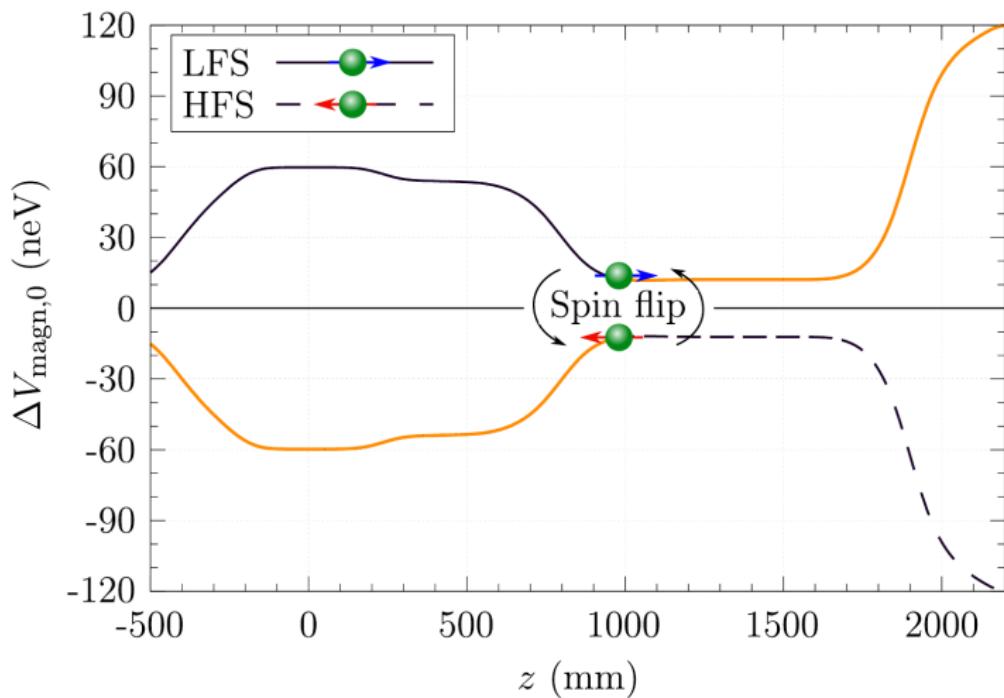
1. Fill UCN into  $\tau$ SPECT Magnet from the left
  - Polarization due to high Magnetic Field, SF on
  - Simultaneously: Intensity Monitoring (non-trappable UCN)
2. Remove SF from storage region
3. Detector to cleaning position and back
4. Wait ...

# Measurement Procedure

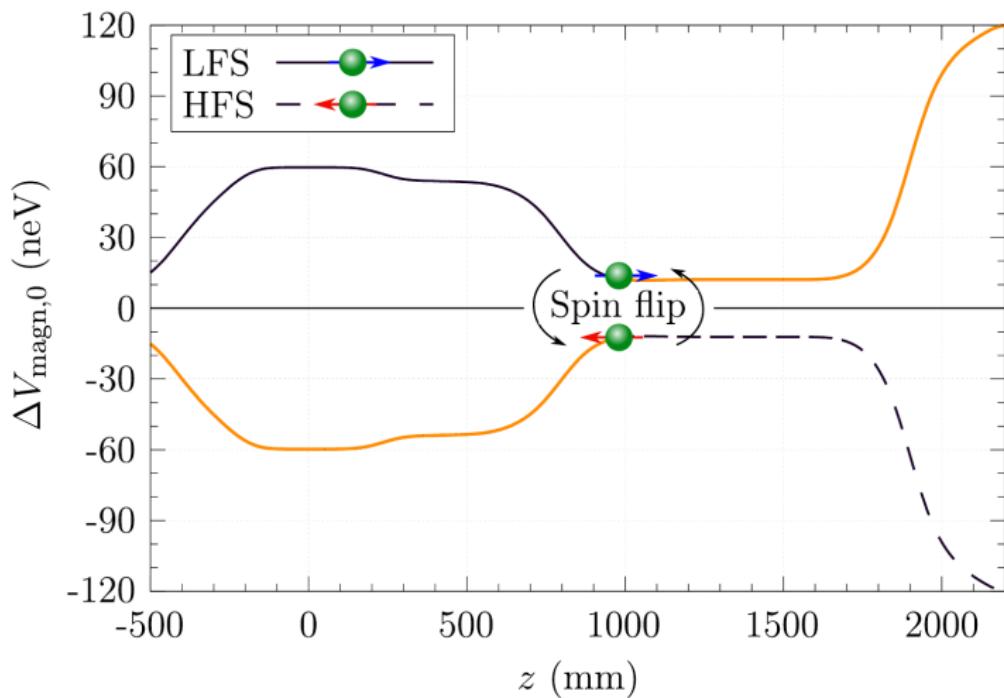


1. Fill UCN into  $\tau$ SPECT Magnet from the left
  - Polarization due to high Magnetic Field, SF on
  - Simultaneously: Intensity Monitoring (non-trappable UCN)
2. Remove SF from storage region
3. Detector to cleaning position and back
4. Wait ...
5. Count UCN

## spin-flip loading

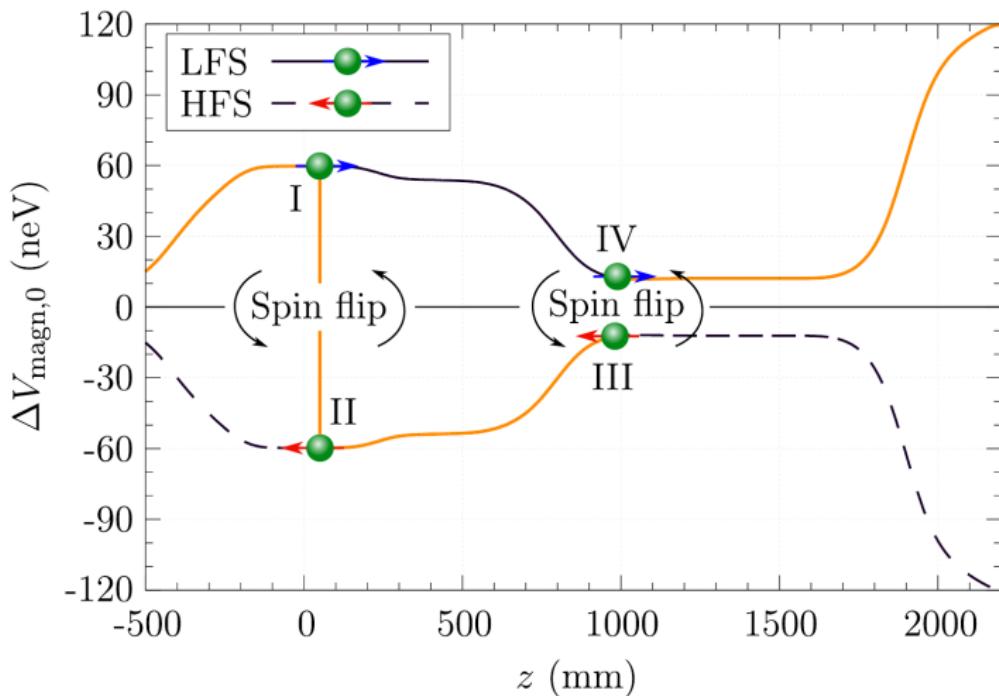


## spin-flip loading

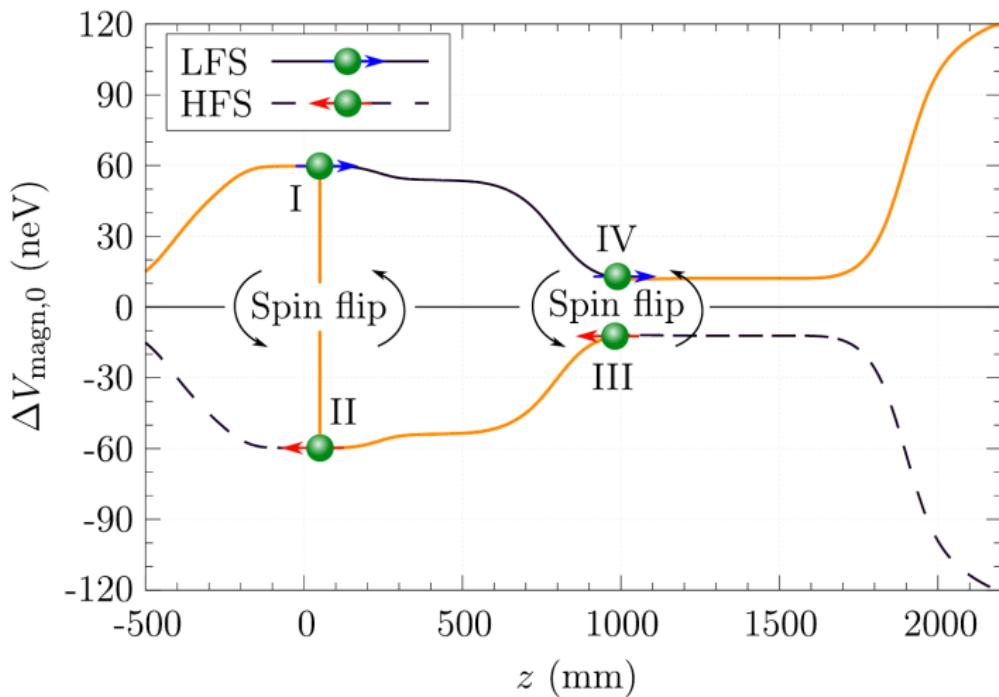


$$E_a = 18.9 \text{ neV}$$

## Double spin-flip loading



## Double spin-flip loading

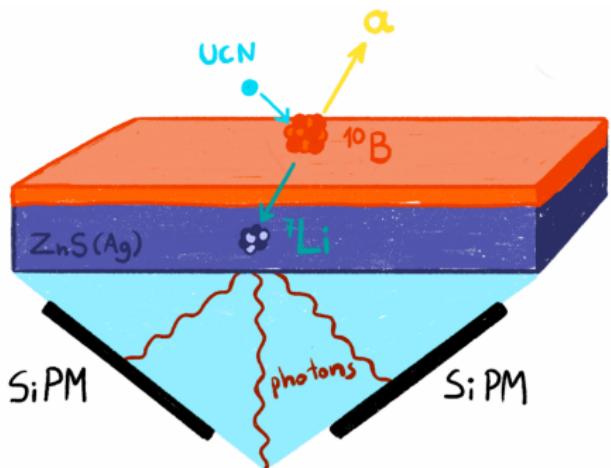


# UCN Intensity Monitor

In order to know  $N(t=0)$ , we need an in-beamline monitor.

# UCN Intensity Monitor

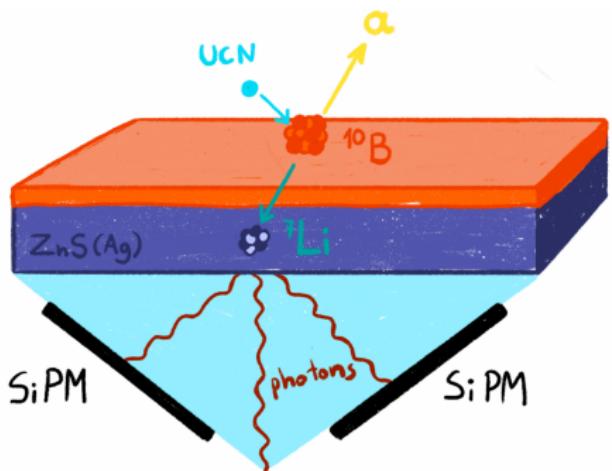
In order to know  $N(t=0)$ , we need an in-beamline monitor.



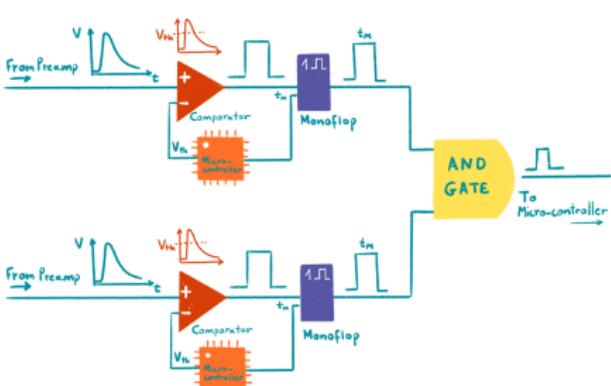
Illustrations: D. Kanta

# UCN Intensity Monitor

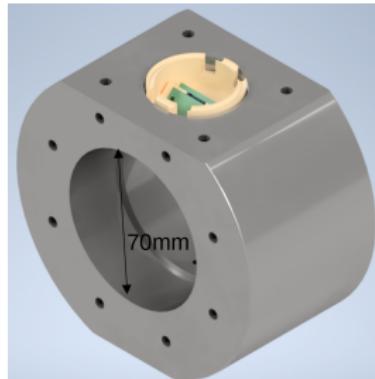
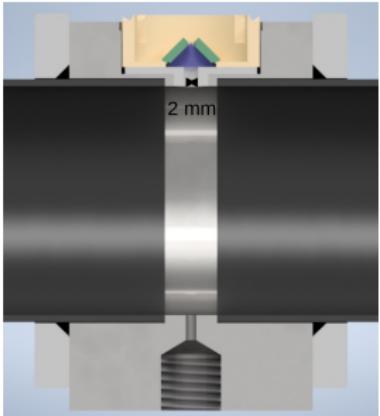
In order to know  $N(t=0)$ , we need an in-beamline monitor.



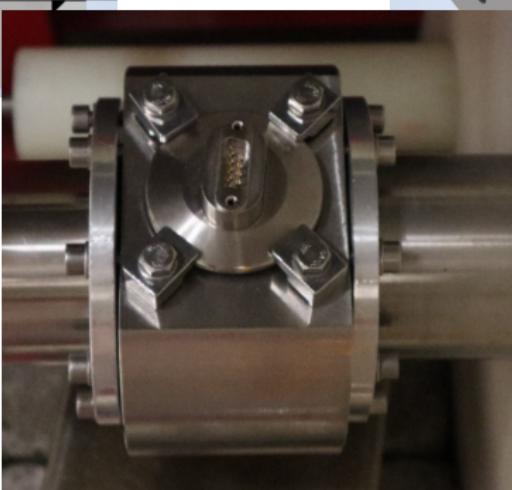
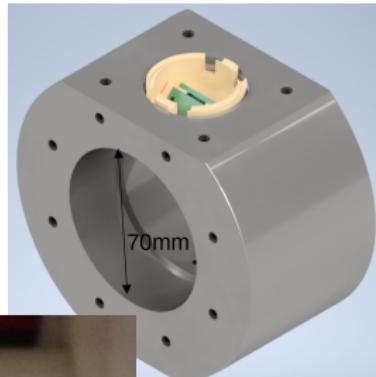
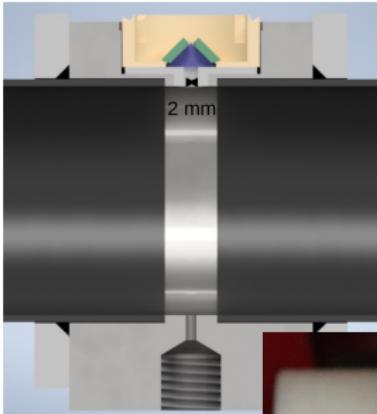
Illustrations: D. Kanta



# UCN Intensity Monitor

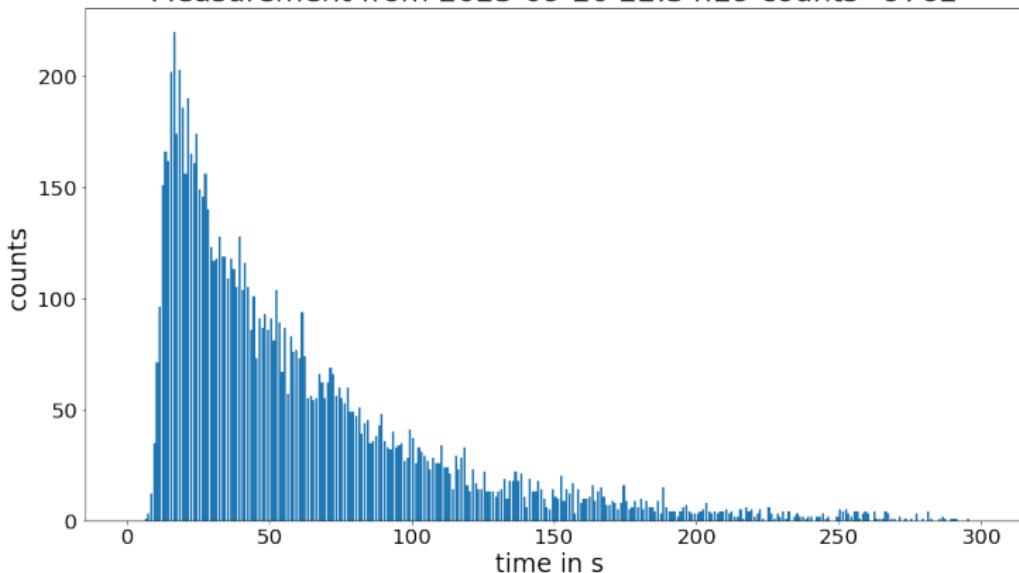


# UCN Intensity Monitor



# UCN Intensity Monitor

Measurement from 2023-09-20 22:34:29 counts=9762



- Detector and readout works!
- Second generation electronics for more robustness and 24/7 operation in development

# Systematics

- Gaps:
- Wall losses:

# Systematics

- Gaps: → 0 ✓
- Wall losses: → 0 ✓
- Depolarisation:

# Systematics

- Gaps: → 0 ✓
- Wall losses: → 0 ✓
- Depolarisation: << 0.1 s ✓
- Rest gas interactions:

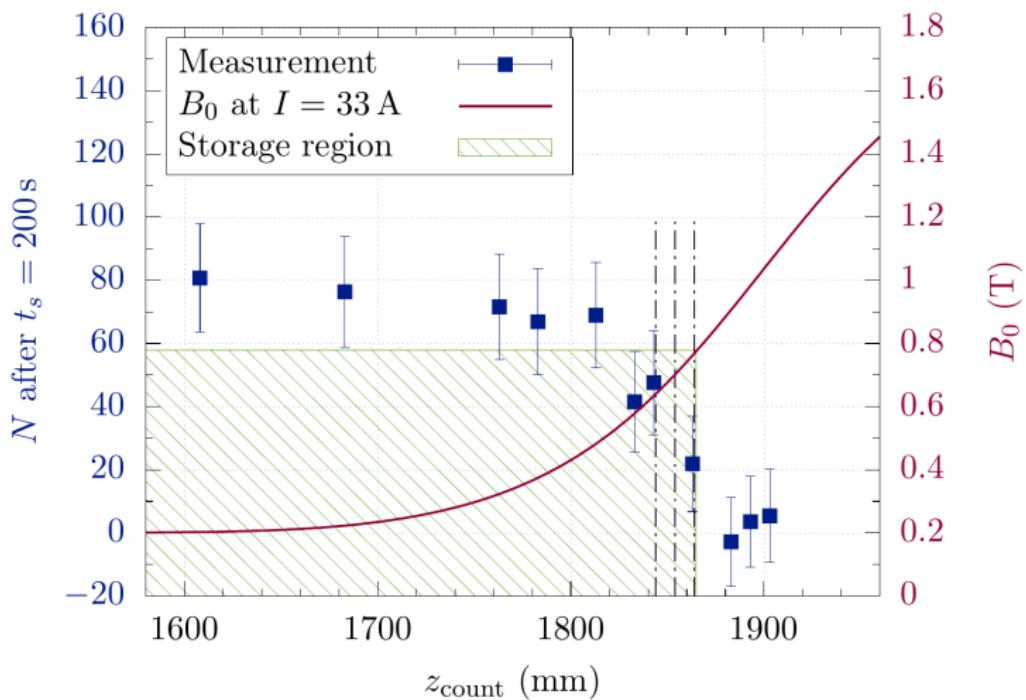
# Systematics

- Gaps: → 0 ✓
- Wall losses: → 0 ✓
- Depolarisation: << 0.1 s ✓
- Rest gas interactions:  $\lesssim 0.1$  s ✓
- Microphonic heating:
- Marginally trapped neutrons:

# Systematics

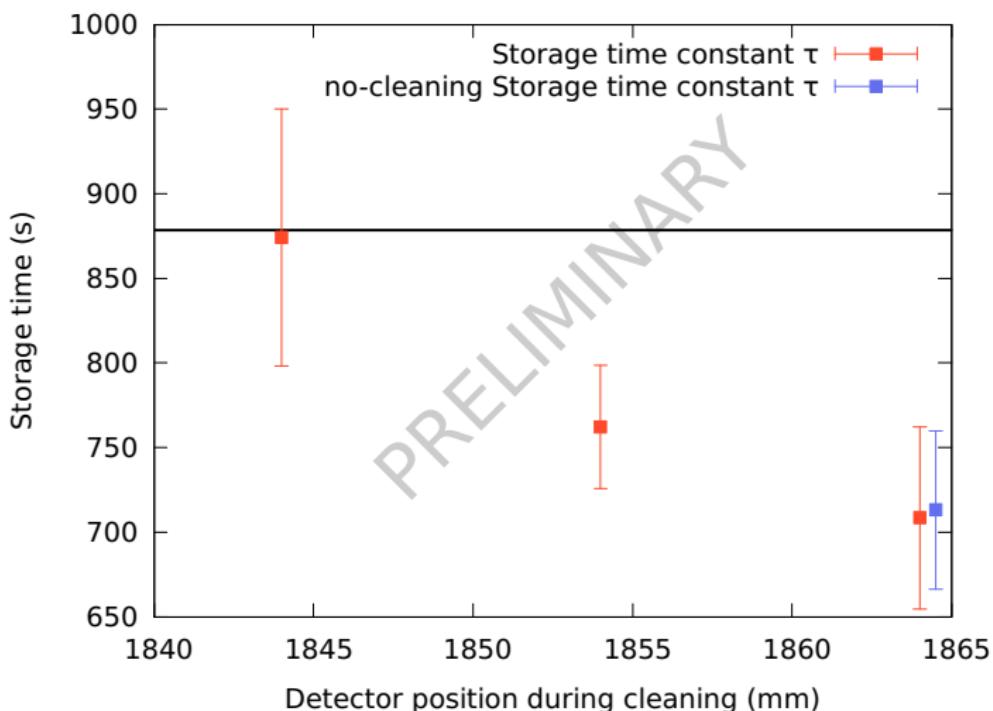
- Gaps: → 0 ✓
- Wall losses: → 0 ✓
- Depolarisation: << 0.1 s ✓
- Rest gas interactions:  $\lesssim 0.1$  s ✓
- Microphonic heating: Has not been observed, measure. ✓
- Marginally trapped neutrons: Spectrum cleaning necessary! ✓

## Countermeasures



K. Ross

# Countermeasures



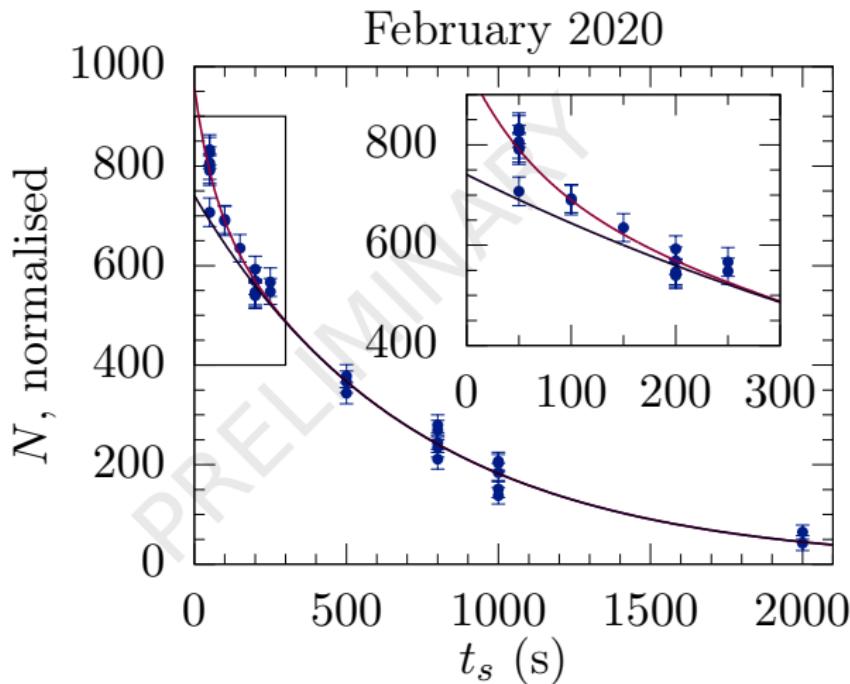
# Systematics Control

- Marginally trapped neutrons:
  - Clean spectrum with active detector before  $t = 0$
  - Demonstrated to work
  - 2 parameters: position and duration
  - Too aggressive cleaning → lower statistics
  - Introduce asymmetry:  $\tau$ SPECT at a small tilt angle

# Systematics Control

- Marginally trapped neutrons:
  - Clean spectrum with active detector before  $t = 0$
  - Demonstrated to work
  - 2 parameters: position and duration
  - Too aggressive cleaning → lower statistics
  - Introduce asymmetry:  $\tau$ SPECT at a small tilt angle
- Microphonic heating:
  - Microphonic heating has not been observed
  - Can be measured via change in arrival time
  - Can be measured by "cleaning" again!
  - New  $\tau$ SPECT timing controller can do that.

## Without Energy Spectrum Cleaning

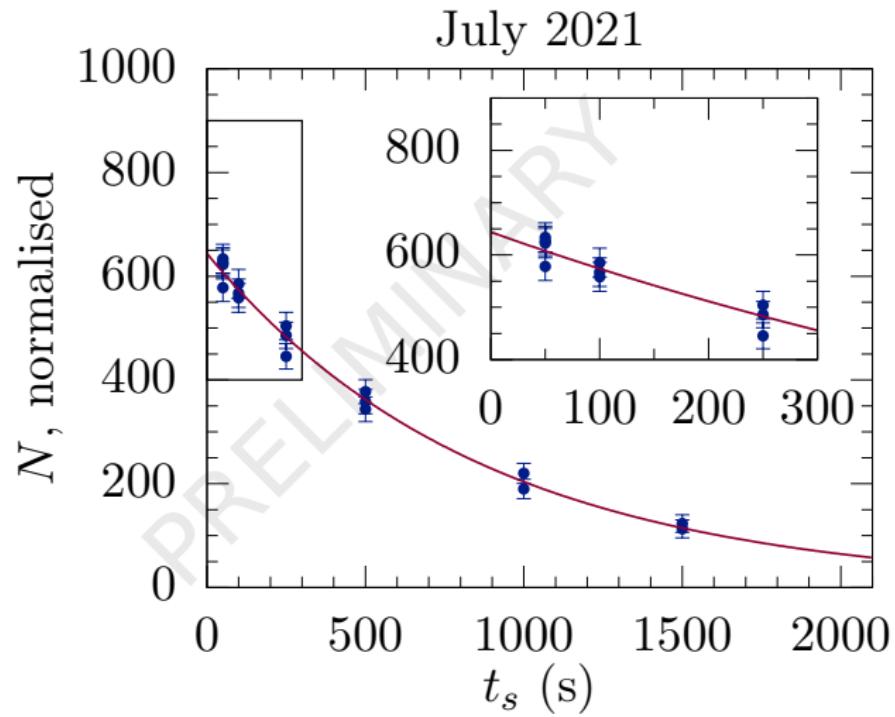


Decay times:  
Fast:  
 $\tau = 64.5$  s  
Slow:  
 $\tau = 740(47)$  s

$$\chi^2 = 1.6$$

K. U. Roß

## With Energy Spectrum Cleaning



Decay times:

$$\tau = 869(29) \text{ s}$$

$$\chi^2 = 0.6$$

K. U. Roß

# $\tau$ SPECT at PSI

UCN CAM 3 Thu Jul 20 05:44:10 2023



# $\tau$ SPECT at PSI

UCN CAM 3 Wed Jul 26 10:22:54 2023



# $\tau$ SPECT at PSI

UCN CAM 3 Sat Aug 19 15:00:27 2023



# Status

- $\tau$ SPECT has been fully commissioned at TRIGA Mainz
- Move and setup to PSI are being concluded
- First neutrons stored in the trap in late 2023!
- $\tau$ SPECT has been approved by PSI's beam time committee in Jan 2024!

# Status

- $\tau$ SPECT has been fully commissioned at TRIGA Mainz
- Move and setup to PSI are being concluded
- First neutrons stored in the trap in late 2023!
- $\tau$ SPECT has been approved by PSI's beam time committee in Jan 2024!

Goal: Show statistical reach and systematics control for a physics run aiming for a precision of 0.1 s in the next years.

# Team



+ W. Heil & P. Blümller

## Team



J. Auler<sup>1</sup>, P. Blümller<sup>1</sup>, M. Engler<sup>2</sup>, M. Fertl<sup>1</sup>, K. Franz<sup>2</sup>, W. Heil<sup>1</sup>,  
S. Kaufmann<sup>2</sup>, N. Pfeifer<sup>1</sup>, D. Ries<sup>3</sup>, N. Yazdandoost<sup>2</sup>

<sup>1</sup> Institute of Physics, Johannes Gutenberg University Mainz, Germany

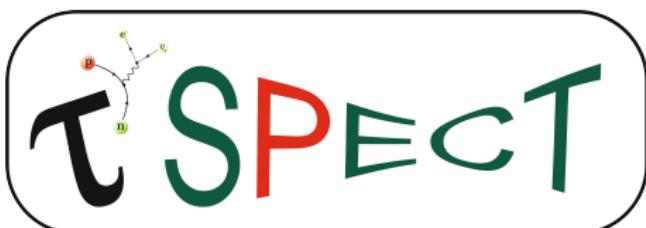
<sup>2</sup> Institute of Nuclear Chemistry, Johannes Gutenberg University Mainz, Germany

<sup>3</sup> Paul Scherrer Institute, Villigen, Switzerland



Supported by the Cluster of Excellence "Precision Physics, Fundamental Interactions, and Structure of Matter" (PRISMA+ EXC 2118/1) funded by the German Research Foundation within the German Excellence Strategy (Project ID 39083149)

## Team



J. Auler<sup>1</sup>, P. Blümller<sup>1</sup>, M. Engler<sup>2</sup>, M. Fertl<sup>1</sup>, K. Franz<sup>2</sup>, W. Heil<sup>1</sup>,  
S. Kaufmann<sup>2</sup>, N. Pfeifer<sup>1</sup>, D. Ries<sup>3</sup>, N. Yazdandoost<sup>2</sup>

<sup>1</sup> Institute of Physics, Johannes Gutenberg University Mainz, Germany

<sup>2</sup> Institute of Nuclear Chemistry, Johannes Gutenberg University Mainz, Germany

<sup>3</sup> Paul Scherrer Institute, Villigen, Switzerland



Supported by the Cluster of Excellence "Precision Physics, Fundamental Interactions, and Structure of Matter" (PRISMA+ EXC 2118/1) funded by the German Research Foundation within the German Excellence Strategy (Project ID 39083149)

Thank you for your attention!